

# ElectroMagnetic (EM) field preliminaries to the observation and interpretation of Cosmic Microwave Background (CMB) radiation

by

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## Synopsis:

- Introduction CMB radiation
- Maxwell - Lorentz ElectroMagnetic field equations for CMB radiation
- EM field and source quantities in CMB radiation
- EM moving particle source quantities in CMB radiation
- Configuration for analyzing CMB radiation
- Far-source approximation in CMB radiation
- Spectral representation of the CMB radiation field
- Observation with the Cosmic Background Imager (CBI)
- State of polarization of the observed field

- **CMB radiation** (equivalent black-body radiation temperature  $T = 2.725$  K, spectral intensity  $I(\nu)$  peaked at  $\nu = 160.2$  GHz) is conjectured to be<sup>[1]</sup>
  - **ElectroMagnetic (EM) radiation** generated by accelerated motion of electrically charged and/or electrically polarized and/or (spin)magnetized particles (Lorentz theory of electrons, 1905), present in ranges of the universe with very low number density of particles
- **Radiation damping** is ascribed to **Thompson scattering**: EM field incident on particle  $\rightarrow$  EM force on particle  $\rightarrow$  acceleration of particle  $\rightarrow$  (Thompson scattered) EM field radiated by particle:
  - ***E*-mode damping**: influence of electric force on particle predominates
  - ***B*-mode damping**: influence of magnetic force (Lorentz force) on particle predominates (usually negligibly small)

- **CMB radiation** is observed and interpreted in **4D space-time** ( $\mathbb{R}^3 \times \mathbb{R}$ ) with
  - **3D Euclidean space**  $\mathbb{R}^3$ 
    - orthonormal Cartesian reference frame  $\{\mathcal{O}, \mathbf{i}_1, \mathbf{i}_2, \mathbf{i}_3\}$
    - position vector  $\mathbf{x} = x_1 \mathbf{i}_1 + x_2 \mathbf{i}_2 + x_3 \mathbf{i}_3 \in \mathbb{R}^3$
  - **1D time**  $t \in \mathbb{R}$
- **Notation:**
  - $\partial_m$  = differentiation with respect to  $x_m$  ( $m = 1, 2, 3$ )
  - $\partial_t$  = differentiation with respect to  $t$

- **Maxwell - Lorentz EM field equations for CMB radiation:**

$$\begin{aligned}
 & \bullet \begin{bmatrix} 0 & -\partial_3 & \partial_2 \\ \partial_3 & 0 & -\partial_1 \\ -\partial_2 & \partial_1 & 0 \end{bmatrix} \begin{bmatrix} H_1 \\ H_2 \\ H_3 \end{bmatrix} - \epsilon_0 \partial_t \begin{bmatrix} E_1 \\ E_2 \\ E_3 \end{bmatrix} = \begin{bmatrix} J_1 \\ J_2 \\ J_3 \end{bmatrix} + \partial_t \begin{bmatrix} P_1 \\ P_2 \\ P_3 \end{bmatrix} \\
 & \bullet \begin{bmatrix} 0 & -\partial_3 & \partial_2 \\ \partial_3 & 0 & -\partial_1 \\ -\partial_2 & \partial_1 & 0 \end{bmatrix} \begin{bmatrix} E_1 \\ E_2 \\ E_3 \end{bmatrix} + \mu_0 \partial_t \begin{bmatrix} H_1 \\ H_2 \\ H_3 \end{bmatrix} = -\mu_0 \partial_t \begin{bmatrix} M_1 \\ M_2 \\ M_3 \end{bmatrix}
 \end{aligned}$$

- $\{E, H\}$  = radiated EM field
- $\{J, P, M\}$  = moving sources generating the field
- $\mu_0$  = (magnetic) permeability of vacuum [ $\mu_0 = 4\pi * 10^{-7}$  H/m ((SI))]
- $\epsilon_0$  = (electric) permittivity of vacuum [ $\epsilon_0 = (\mu_0 c_0^2)^{-1}$  F/m (SI)]
- $c_0$  = electromagnetic wave speed in vacuum [ $c_0 = 299792458$  m/s (SI)]

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Maxwell (1873) - Lorentz (1905) ElectroMagnetic field equations for CMB radiation

- **EM field quantities in CMB radiation:**

- $E(\mathbf{x}, t)$  = electric field strength (V/m)

- $H(\mathbf{x}, t)$  = magnetic field strength (A/m)

- **EM source quantities in CMB radiation:**

- $\mathbf{J}(\mathbf{x}, t)$  = volume density of electric convection current (A/m<sup>2</sup>)

- $\mathbf{P}(\mathbf{x}, t)$  = volume density of electric polarization (C/m<sup>2</sup>)

- $\mathbf{M}(\mathbf{x}, t)$  = volume density of (spin) magnetization (A/m)

- EM moving particle **source** quantities in CMB radiation:<sup>[2]</sup>
  - $\mathbf{x}_0(t)$  = position of particle at time  $t$
  - $\mathbf{v}_0(t)$  = velocity of particle at time  $t$  [ $\mathbf{v}_0(t) = d_t \mathbf{x}_0(t)$ ]
  - $\delta^3(\mathbf{x})$  = 3D Dirac delta distribution operative at  $\mathbf{x} = \mathbf{0}$ 
    - $\delta^3(\mathbf{x}) = 0$  for  $\mathbf{x} \neq \mathbf{0}$
    - $\int_{\mathbf{x} \in \mathbb{R}^3} \delta^3[\mathbf{x} - \mathbf{x}_0(t)] f(\mathbf{x}, t) dV = f[\mathbf{x}_0(t), t]$  ("sifting property")
  - $\mathbf{J}(\mathbf{x}, t) = q(t) \mathbf{v}_0(t) \delta^3[\mathbf{x} - \mathbf{x}_0(t)]$ 
    - $q(t)$  = electric charge (C)
  - $\mathbf{P}(\mathbf{x}, t) = \mathbf{p}(t) \delta^3[\mathbf{x} - \mathbf{x}_0(t)]$ 
    - $\mathbf{p}(t)$  = electric dipole moment (C·m)
  - $\mathbf{M}(\mathbf{x}, t) = \mathbf{m}(t) \delta^3[\mathbf{x} - \mathbf{x}_0(t)]$ 
    - $\mathbf{m}(t)$  = magnetic spin moment (A·m<sup>2</sup>)



- Note that

- $$[\partial_1 \ \partial_2 \ \partial_3] \begin{bmatrix} 0 & \partial_3 & -\partial_2 \\ -\partial_3 & 0 & \partial_1 \\ \partial_2 & -\partial_1 & 0 \end{bmatrix} = [0 \ 0 \ 0] \quad \text{implies:}$$

- EM compatibility relations

- $$[\partial_1 \ \partial_2 \ \partial_3] \begin{bmatrix} -\epsilon_0 \partial_t E_1 \\ -\epsilon_0 \partial_t E_2 \\ -\epsilon_0 \partial_t E_3 \end{bmatrix} = [\partial_1 \ \partial_2 \ \partial_3] \begin{bmatrix} J_1 + \partial_t P_1 \\ J_2 + \partial_t P_2 \\ J_3 + \partial_t P_3 \end{bmatrix}$$

- $$[\partial_1 \ \partial_2 \ \partial_3] \begin{bmatrix} -\mu_0 \partial_t H_1 \\ -\mu_0 \partial_t H_2 \\ -\mu_0 \partial_t H_3 \end{bmatrix} = [\partial_1 \ \partial_2 \ \partial_3] \begin{bmatrix} -\mu_0 \partial_t M_1 \\ -\mu_0 \partial_t M_2 \\ -\mu_0 \partial_t M_3 \end{bmatrix}$$

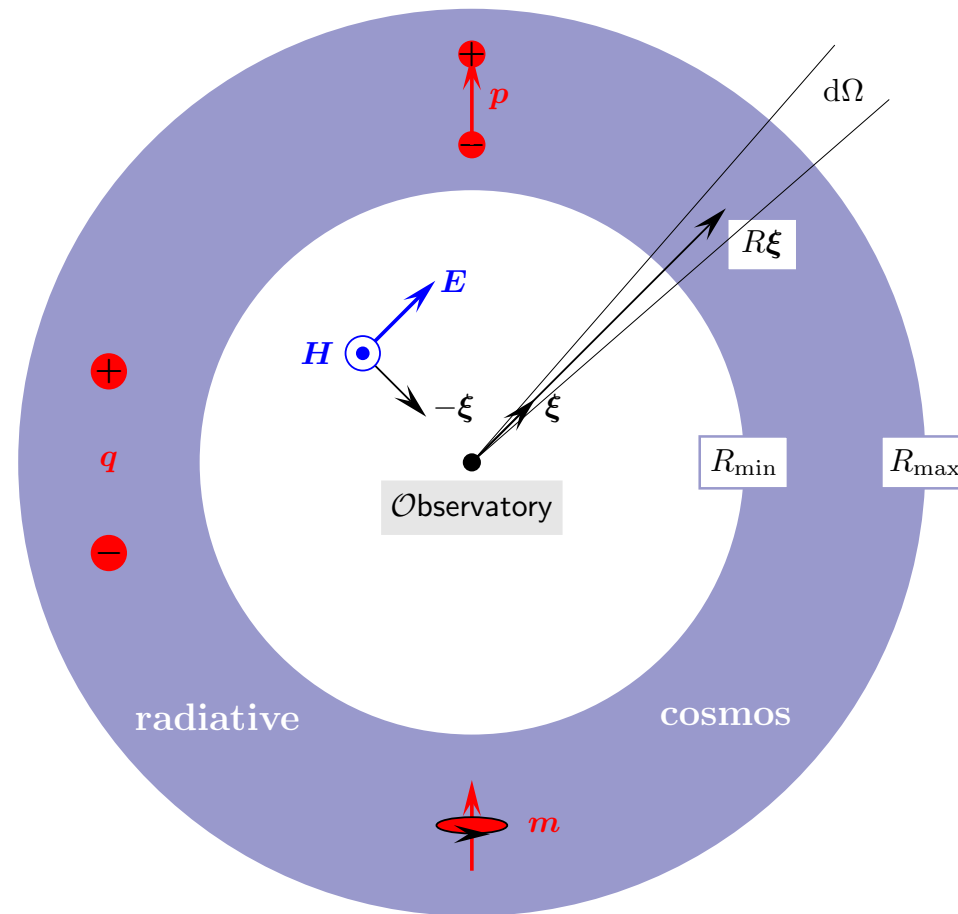


Fig. 1. Radiative cosmos with moving electric charges, electric dipoles and magnetic (spin) dipoles, and observation coordinates for CMB radiation.

- Maxwell field calculation/computation in operator form:

- $\mathbf{G}: \{J, P, M\}(\mathbf{x}', t') \xrightarrow{\mathbf{G}} \{E, H\}(\mathbf{x}, t)$

- $\{J, P, M\}(\mathbf{x}', t') =$  sources

- $\{E, H\}(\mathbf{x}, t) =$  field

- $\mathbf{G}(\mathbf{x}, t; \mathbf{x}', t') =$  Maxwell field propagator

- Far-source approximation: •  $c_0/|\mathbf{x}' - \mathbf{x}| \ll \partial_t \implies \partial_m \simeq (\xi_m/c_0)\partial_t$

- $\mathbf{G}^{\text{far}}: \{J, P, M\}(\mathbf{x}', t') \xrightarrow{\mathbf{G}^{\text{far}}} \{E, H\}^{\text{far}}(\mathbf{x}, t)$

- $\{J, P, M\}(\mathbf{x}', t') =$  sources

- $\{E, H\}^{\text{far}}(\mathbf{x}, t) =$  field in far-source approximation

- $\mathbf{G}^{\text{far}}(\mathbf{x}, t; \mathbf{x}', t') =$  far-source Maxwell field propagator

- $\xi_m = \frac{\mathbf{x}' - \mathbf{x}}{|\mathbf{x}' - \mathbf{x}|} \rightarrow \frac{\mathbf{x}'}{|\mathbf{x}'|}$  as  $|\mathbf{x}'| \rightarrow \infty$

- The CMB observed field (far-source approximation):

- $\{E_m, H_m\}(\mathbf{x}, t) = \frac{1}{4\pi} \int_{\xi \in \Omega} \{E_m^\infty, H_m^\infty\}(\xi, t) d\Omega \quad (m = 1, 2, 3)$

- $$\begin{bmatrix} E_1^\infty \\ E_2^\infty \\ E_3^\infty \end{bmatrix} = \mu_0 \begin{bmatrix} \xi_1 \xi_1 - 1 & \xi_1 \xi_2 & \xi_1 \xi_3 \\ \xi_2 \xi_1 & \xi_2 \xi_2 - 1 & \xi_2 \xi_3 \\ \xi_3 \xi_1 & \xi_3 \xi_2 & \xi_3 \xi_3 - 1 \end{bmatrix} \begin{bmatrix} \partial_t A_1 \\ \partial_t A_2 \\ \partial_t A_3 \end{bmatrix} + \frac{1}{c_0} \begin{bmatrix} 0 & -\xi_3 & \xi_2 \\ \xi_3 & 0 & -\xi_1 \\ -\xi_2 & \xi_1 & 0 \end{bmatrix} \begin{bmatrix} \partial_t F_1 \\ \partial_t F_2 \\ \partial_t F_3 \end{bmatrix}$$
- $$\begin{bmatrix} H_1^\infty \\ H_2^\infty \\ H_3^\infty \end{bmatrix} = \epsilon_0 \begin{bmatrix} \xi_1 \xi_1 - 1 & \xi_1 \xi_2 & \xi_1 \xi_3 \\ \xi_2 \xi_1 & \xi_2 \xi_2 - 1 & \xi_2 \xi_3 \\ \xi_3 \xi_1 & \xi_3 \xi_2 & \xi_3 \xi_3 - 1 \end{bmatrix} \begin{bmatrix} \partial_t F_1 \\ \partial_t F_2 \\ \partial_t F_3 \end{bmatrix} - \frac{1}{c_0} \begin{bmatrix} 0 & -\xi_3 & \xi_2 \\ \xi_3 & 0 & -\xi_1 \\ -\xi_2 & \xi_1 & 0 \end{bmatrix} \begin{bmatrix} \partial_t A_1 \\ \partial_t A_2 \\ \partial_t A_3 \end{bmatrix}$$

- $A_m(\xi, t) = \int_{R=R_{\min}}^{R_{\max}} [J_m(R\xi, t - R/c_0) + \partial_t P_m(R\xi, t - R/c_0)] R dR \quad (m = 1, 2, 3)$

- $F_m(\xi, t) = \int_{R=R_{\min}}^{R_{\max}} \partial_t M_m(R\xi, t - R/c_0) R dR \quad (m = 1, 2, 3)$

- Far-source relationships EM field equations [ $\partial_m \simeq (\xi_m/c_0)\partial_t$ ]:

$$\bullet \begin{bmatrix} 0 & -\xi_3 & \xi_2 \\ \xi_3 & 0 & -\xi_1 \\ -\xi_2 & \xi_1 & 0 \end{bmatrix} \begin{bmatrix} H_1^\infty \\ H_2^\infty \\ H_3^\infty \end{bmatrix} - c_0\epsilon_0 \begin{bmatrix} E_1^\infty \\ E_2^\infty \\ E_3^\infty \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \end{bmatrix}$$

$$\bullet \begin{bmatrix} 0 & -\xi_3 & \xi_2 \\ \xi_3 & 0 & -\xi_1 \\ -\xi_2 & \xi_1 & 0 \end{bmatrix} \begin{bmatrix} E_1^\infty \\ E_2^\infty \\ E_3^\infty \end{bmatrix} + c_0\mu_0 \begin{bmatrix} H_1^\infty \\ H_2^\infty \\ H_3^\infty \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ 0 \end{bmatrix}$$

- Far-source relationships EM compatibility relations [ $\partial_m \simeq (\xi_m/c_0)\partial_t$ ]:

$$\bullet \begin{bmatrix} \xi_1 & \xi_2 & \xi_3 \end{bmatrix} \begin{bmatrix} E_1^\infty \\ E_2^\infty \\ E_3^\infty \end{bmatrix} = 0 \quad \bullet \begin{bmatrix} \xi_1 & \xi_2 & \xi_3 \end{bmatrix} \begin{bmatrix} H_1^\infty \\ H_2^\infty \\ H_3^\infty \end{bmatrix} = 0$$

- **Spectral representation of electric and magnetic field strength (Fourier representation):**

- $$\begin{aligned} \{ \mathbf{E}, \mathbf{H} \}(\mathbf{x}, t) &= \frac{1}{2\pi} \int_{\omega=-\infty}^{\infty} \exp(j\omega t) \{ \hat{\mathbf{E}}, \hat{\mathbf{H}} \}(\mathbf{x}, j\omega) d\omega \\ &\quad (\omega = \text{angular frequency}) \\ &= \int_{\nu=-\infty}^{\infty} \exp(j2\pi\nu t) \{ \hat{\mathbf{E}}, \hat{\mathbf{H}} \}(\mathbf{x}, j2\pi\nu) d\nu \\ &\quad (\nu = \text{frequency} = \omega/2\pi) \end{aligned}$$

- **Fourier theorem:**

$$\{ \hat{\mathbf{E}}, \hat{\mathbf{H}} \}(\mathbf{x}, j\omega) = \int_{t=-\infty}^{\infty} \exp(-j\omega t) \{ \mathbf{E}, \mathbf{H} \}(\mathbf{x}, t) dt$$

- **Since the field vectors are real-valued, also:**

$$\{ \mathbf{E}, \mathbf{H} \}(\mathbf{x}, t) = \frac{1}{\pi} \int_{\omega=0}^{\infty} \text{Re} \left[ \exp(j\omega t) \{ \hat{\mathbf{E}}, \hat{\mathbf{H}} \}(\mathbf{x}, j\omega) \right] d\omega$$

- For any set of components of the electric and the magnetic field strength (Parseval's theorem):

$$\begin{aligned} \bullet \int_{t=-\infty}^{\infty} E(\mathbf{x}, t)H(\mathbf{x}, t)dt &= \frac{1}{2\pi} \int_{\omega=-\infty}^{\infty} \hat{E}(\mathbf{x}, j\omega)\hat{H}(\mathbf{x}, -j\omega)d\omega \\ &= \int_{\nu=-\infty}^{\infty} \hat{E}(\mathbf{x}, j2\pi\nu)\hat{H}(\mathbf{x}, -j2\pi\nu)d\nu \end{aligned}$$

- $\hat{E}(\mathbf{x}, j\omega)\hat{H}(\mathbf{x}, -j\omega) = \text{spectral power density}$

- Since the field vectors are real-valued, also:

$$\int_{t=-\infty}^{\infty} E(\mathbf{x}, t)H(\mathbf{x}, t)dt = \frac{1}{\pi} \int_{\omega=0}^{\infty} \text{Re} \left[ \hat{E}(\mathbf{x}, j\omega)\hat{H}(\mathbf{x}, -j\omega) \right] d\omega$$

**Notation:** Vectors are denoted by bold-face symbols,  $\cdot$  denotes inner product,

- $|\mathbf{E}|^2 = \mathbf{E} \cdot \mathbf{E}$

- **Each spectral field component:**

- $\mathbf{E}(\mathbf{x}, t) = \text{Re}[\hat{\mathbf{E}}(\mathbf{x}, j\omega)] \cos(\omega t) - \text{Im}[\hat{\mathbf{E}}(\mathbf{x}, j\omega)] \sin(\omega t)$

describes at position  $\mathbf{x}$  in the plane through  $\text{Re}[\hat{\mathbf{E}}(\mathbf{x}, j\omega)]$  and  $\text{Im}[\hat{\mathbf{E}}(\mathbf{x}, j\omega)]$  a **closed curve** in the course of time over any cycle of duration  $2\pi/\omega$ .

- **In:**

- $\mathbf{E}(\mathbf{x}, t) = \mathbf{E}_X \cos(\omega t + \psi) + \mathbf{E}_Y \sin(\omega t + \psi)$

$\{\mathbf{E}_X, \mathbf{E}_Y\}$  can be constructed such that

- $\mathbf{E}_X \cdot \mathbf{E}_Y = 0 \quad (\mathbf{E}_X \perp \mathbf{E}_Y)$



- **Since:**

- $\mathbf{E}_X = \text{Re}(\hat{\mathbf{E}}) \cos(\psi) + \text{Im}(\hat{\mathbf{E}}) \sin(\psi) \quad (\omega t = 0)$

- $\mathbf{E}_Y = \text{Re}(\hat{\mathbf{E}}) \sin(\psi) - \text{Im}(\hat{\mathbf{E}}) \cos(\psi) \quad (\omega t = \pi/2)$

- $\tan(2\psi)$  is found as

- $\tan(2\psi) = \frac{2 \text{Re}(\hat{\mathbf{E}}) \cdot \text{Im}(\hat{\mathbf{E}})}{|\text{Re}(\hat{\mathbf{E}})|^2 - |\text{Im}(\hat{\mathbf{E}})|^2}$

- **while:**

- $\mathbf{E} \cdot \mathbf{E}_X = |\mathbf{E}_X|^2 \cos(\omega t + \psi)$
  - $\mathbf{E} \cdot \mathbf{E}_Y = |\mathbf{E}_Y|^2 \sin(\omega t + \psi)$

- $\implies$

- $\left( \frac{\mathbf{E} \cdot \mathbf{E}_X}{|\mathbf{E}_X|^2} \right)^2 + \left( \frac{\mathbf{E} \cdot \mathbf{E}_Y}{|\mathbf{E}_Y|^2} \right)^2 = 1 \quad \longleftarrow \text{ellipse if } |\mathbf{E}_X|^2 \neq |\mathbf{E}_Y|^2$

- **Semi major axis of ellipse (if •  $|E_X|^2 \neq |E_Y|^2$ ) :**

- $|E_X|^2 = \frac{1}{2} \left( |\text{Re}(\hat{E})|^2 + |\text{Im}(\hat{E})|^2 \right) + \left[ \frac{1}{4} \left( |\text{Re}(\hat{E})|^2 - |\text{Im}(\hat{E})|^2 \right)^2 + \left( \text{Re}(\hat{E}) \cdot \text{Im}(\hat{E}) \right)^2 \right]^{1/2}$

- **Semi minor axis of ellipse (if •  $|E_X|^2 \neq |E_Y|^2$ ) :**

- $|E_Y|^2 = \frac{1}{2} \left( |\text{Re}(\hat{E})|^2 + |\text{Im}(\hat{E})|^2 \right) - \left[ \frac{1}{4} \left( |\text{Re}(\hat{E})|^2 - |\text{Im}(\hat{E})|^2 \right)^2 + \left( \text{Re}(\hat{E}) \cdot \text{Im}(\hat{E}) \right)^2 \right]^{1/2}$

**If: •  $|E_X|^2 = |E_Y|^2$  ellipse  $\implies$  circle**

**If: •  $E_X \parallel E_Y$  ellipse  $\implies$  line**

- [1] [http://en.wikipedia.org/wiki/Cosmic\\_microwave\\_background\\_radiation](http://en.wikipedia.org/wiki/Cosmic_microwave_background_radiation)
- [2] De Hoop, A. T., Fields and waves excited by point sources in motion – The general 3D time-domain Doppler effect, *Wave Motion*, 43, 2005, pp. 116-122
- [3] De Hoop, A. T., *Theorie van het elektromagnetische veld*, Delftse Universitaire Pers, 1975

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## References